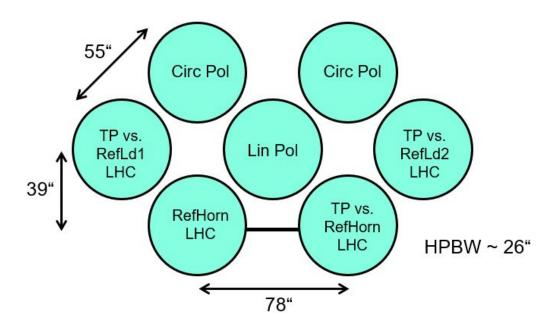
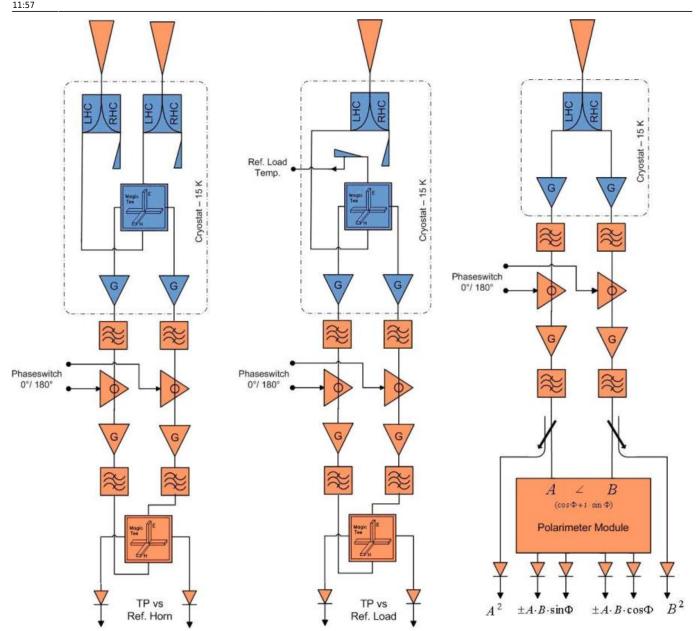
Technical Documentation of the EMBA Receiver 30 -34 GHz (S9mm)

| Туре | HEMT cooled | |
|-----------------|-------------------|---|
| Channels | 12 | (4 via V/F converters) |
| Receiver Noise | 18 K - 24 K | (about 55 K System Noise Temperature) |
| Frequency range | 33 GHz - 50 GHz | |
| Polarization | LHC/RHC or linear | see horn layout |
| Calibration | Noise diode | Transmitter optically coupled |
| IF | none | direct detection principle for continuum only |
| | | |

The receiver is a pseudocorrelation design similar to those used for WMAP / PLANCK that offers suppression of the 1/f noise resulting from gain fluctuations which are inherent to cryogenic InP HEMTs. There will be three polarimetry pixels available, 2 circular, 1 linear polarized. Another two pixels will be total power LCP, each referenced to an internal cold waveguide load whose temperature is accurately measured synchronously with the data from the RF detectors. Finally there is beam switch implemented in hardware with two horns differenced by means of a waveguide magic tee. In figure 2 a schematic sketch of the different types of RF channels that are implemented in the receiver is given. Since the receiver is laid out as a direct detection receiver, no mixing to an IF at a few GHz is carried out, all processing down to the final direct detection step is done at 30 - 34 GHz.



Horn layout of the 7 beam 9mm receiver with beam spacing on sky indicated



Schematic diagram of the different types of RF channels for the 7 beam 9mm receiver

Short description of the receiver (FS on 22.1.2008)

https://eff100mwiki.mpifr-bonn.mpg.de/ - Effelsberg 100m Teleskop

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Last update: 2019/12/13 11:57

