

Mapping polarized continuum emission with the Effelsberg Telescope

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Continuum & polarization mapping: present status

- **Single-horn receivers** in primary focus (1.4 GHz) or in secondary focus (2.64 and 8.35 GHz)
- **Multi-horn receivers** in secondary focus (4.85, 10.45, 32 GHz)
- Linearly polarized signals (X and Y) are transformed into **circularly polarized** signals (Stokes parameters R and L) by a hardware hybrid, turning the phase of one signal by 90°
- A **correlator** generates signals of **linear polarization** in Stokes parameters U and Q
- **4 output signals: R, L, U, Q**
- **But only one broad frequency channel**

Computing Stokes parameters by correlation

(1) Linear input:

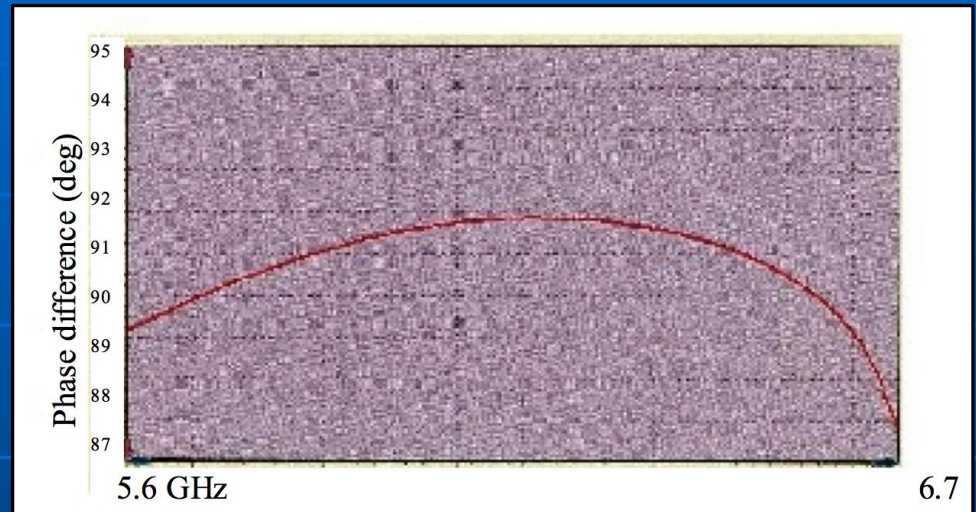
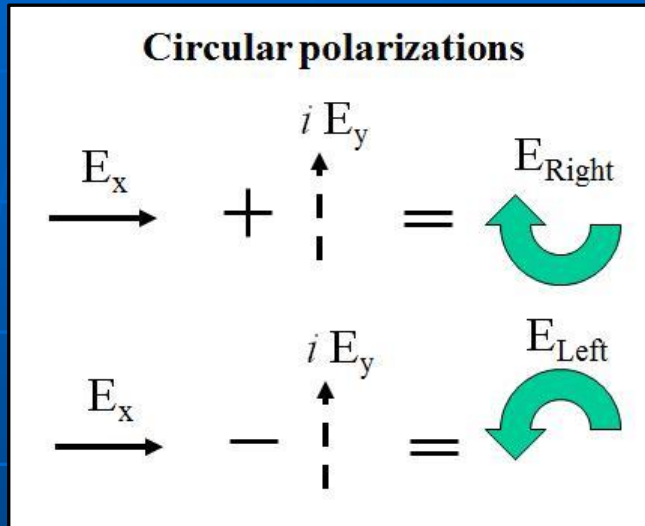
$$\begin{aligned} I &= |E_x|^2 + |E_y|^2, \\ Q &= |E_x|^2 - |E_y|^2, \\ U &= 2\operatorname{Re}(E_x E_y^*), \\ V &= -2\operatorname{Im}(E_x E_y^*), \end{aligned}$$

(2) Circular input:

$$\begin{aligned} I &= |E_l|^2 + |E_r|^2, \\ Q &= 2\operatorname{Re}(E_l^* E_r), \\ U &= -2\operatorname{Im}(E_l^* E_r), \\ V &= |E_l|^2 - |E_r|^2. \end{aligned}$$

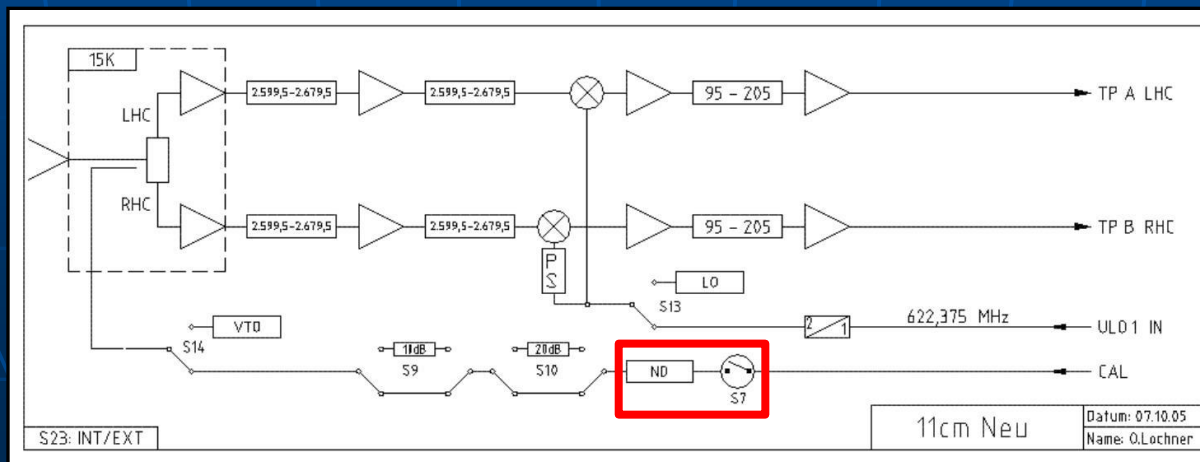
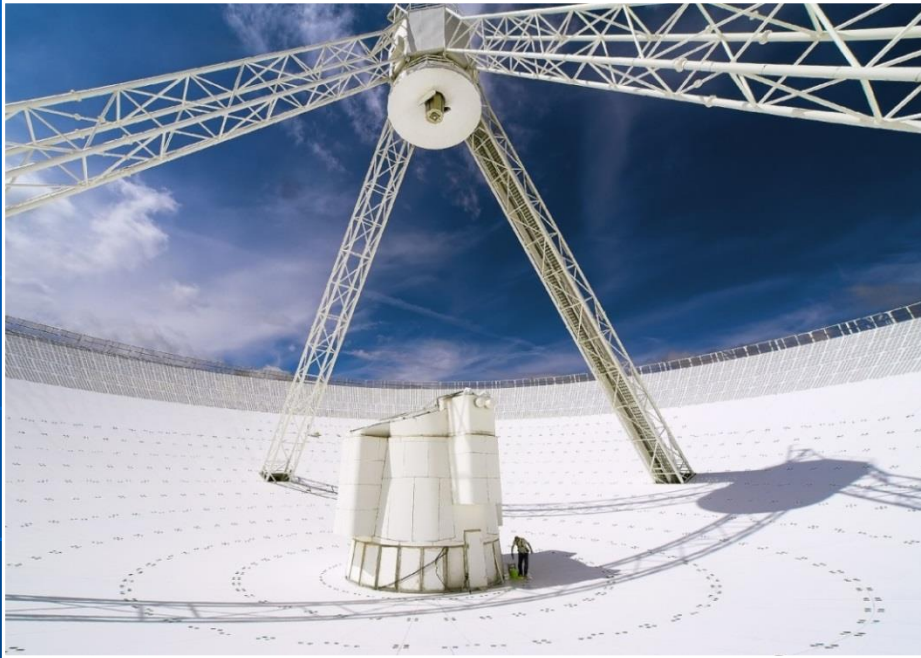
Stable signals in Q and U need circular inputs !

Hybrid



- No perfect phase turning possible (elliptical pol)
- Wide bands ($\Delta\nu/\nu > 2$): hardware hybrid no longer appropriate

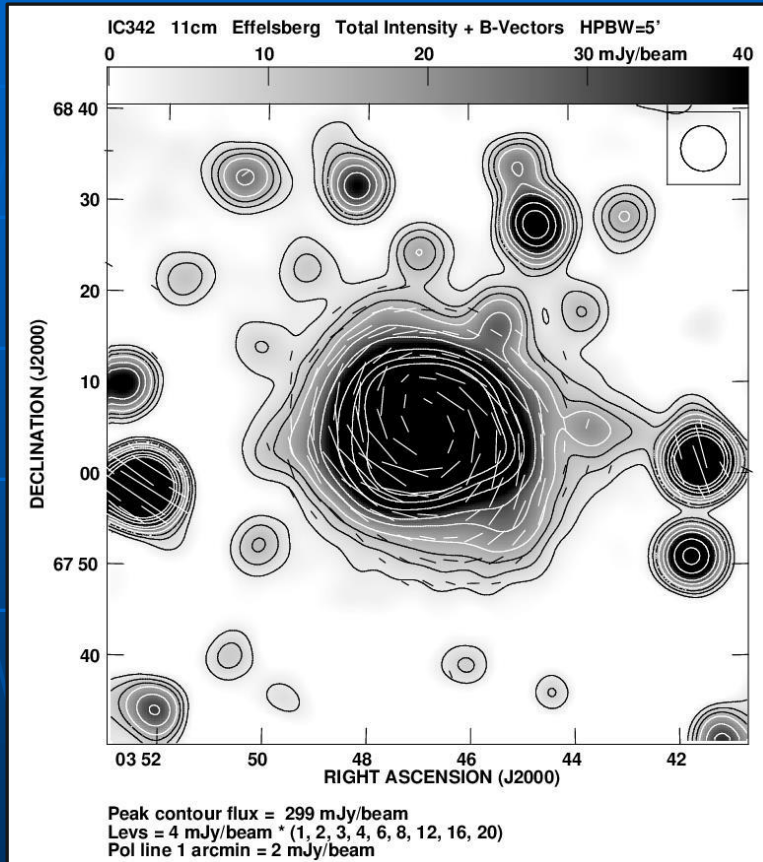
Effelsberg 2.64 GHz (11 cm) SFK single-horn receiver



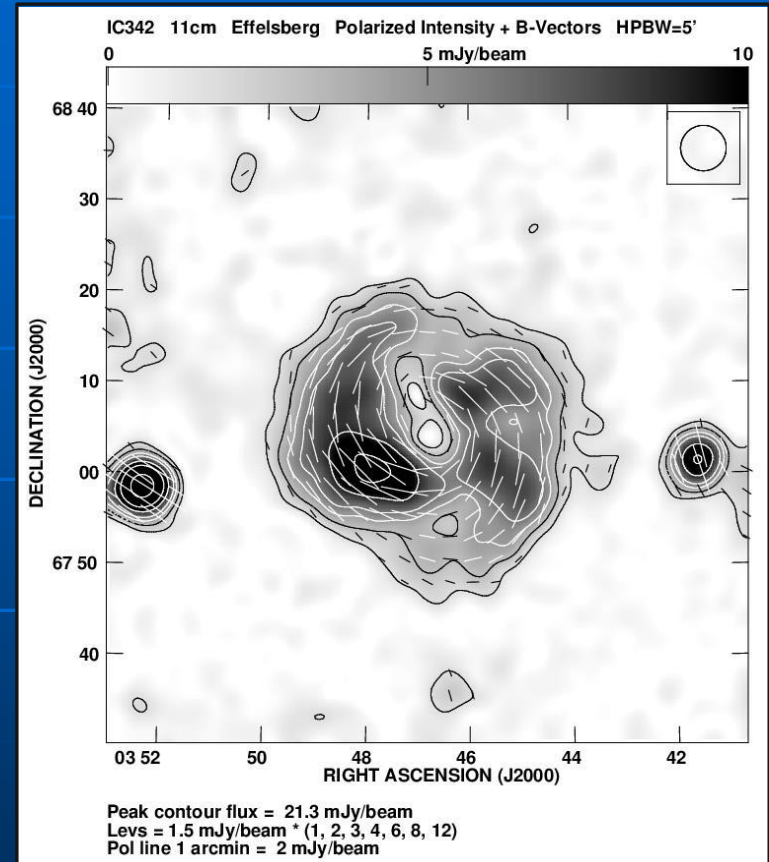
Continuum & polarization mapping: calibration signal

- Noise diode: linearly polarized calibration signal to remove variations in gain and phase
- On/off switching at 32 msec rate
- Signal and calibration signals are separated in 32 msec intervals
- Strong RFI is filtered by comparing adjacent points
- Calibration signal is linearly fitted for each subscan (a single scan in a map) in all 4 signal channels
- Signal/ $\langle \text{cal} \rangle$ is averaged to about 3 pixels per beamwidth

Effelsberg 2.64 GHz (11.1 cm), BW = 80 MHz

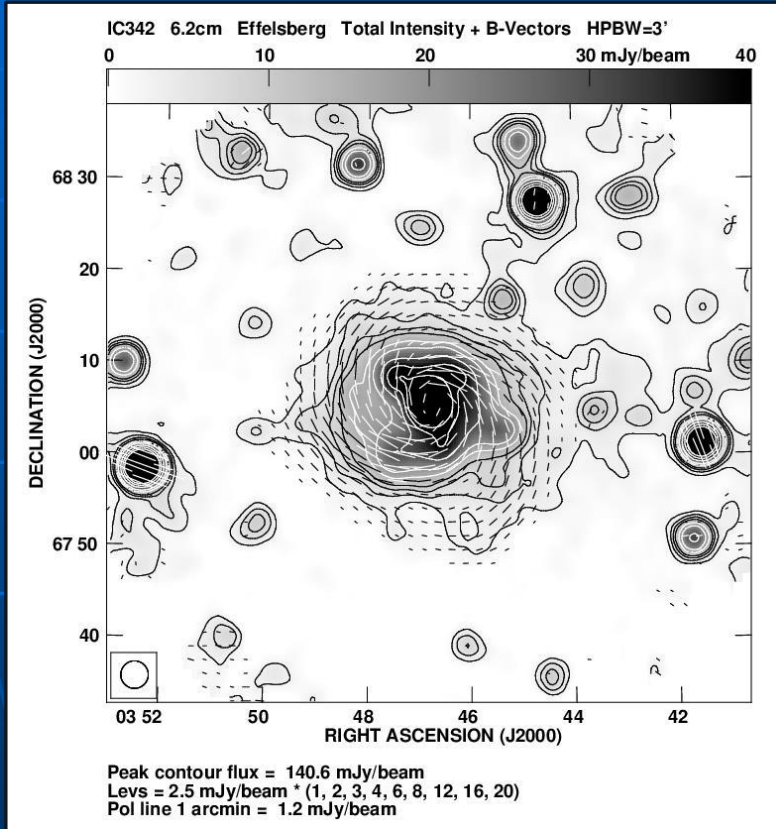


Total intensity:
 $\sigma \approx 0.8 \text{ mJy} / 5' \text{ beam}$

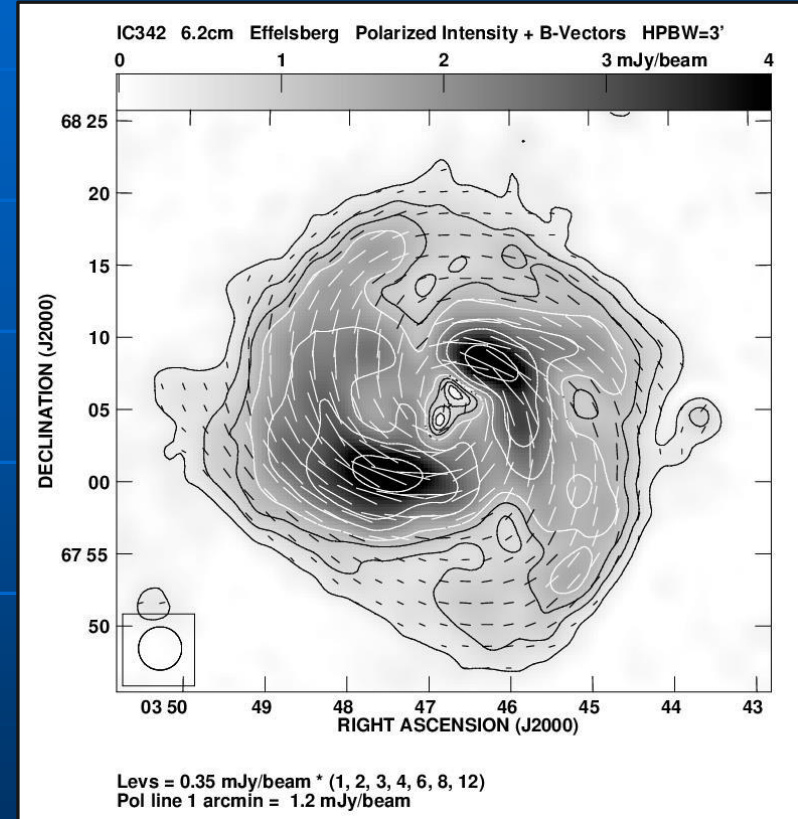


Q & U:
 $\sigma \approx 0.45 \text{ mJy} / 5' \text{ beam}$

Effelsberg 4.85 GHz (6.3 cm), BW = 500 MHz

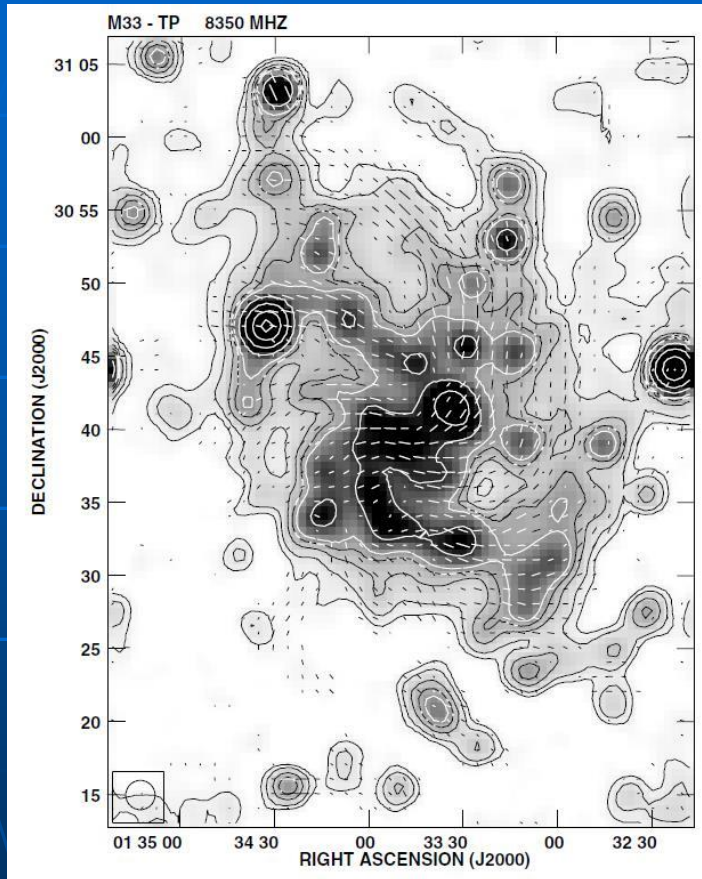


Total intensity:
 $\sigma \approx 0.5 \text{ mJy} / 3' \text{ beam}$

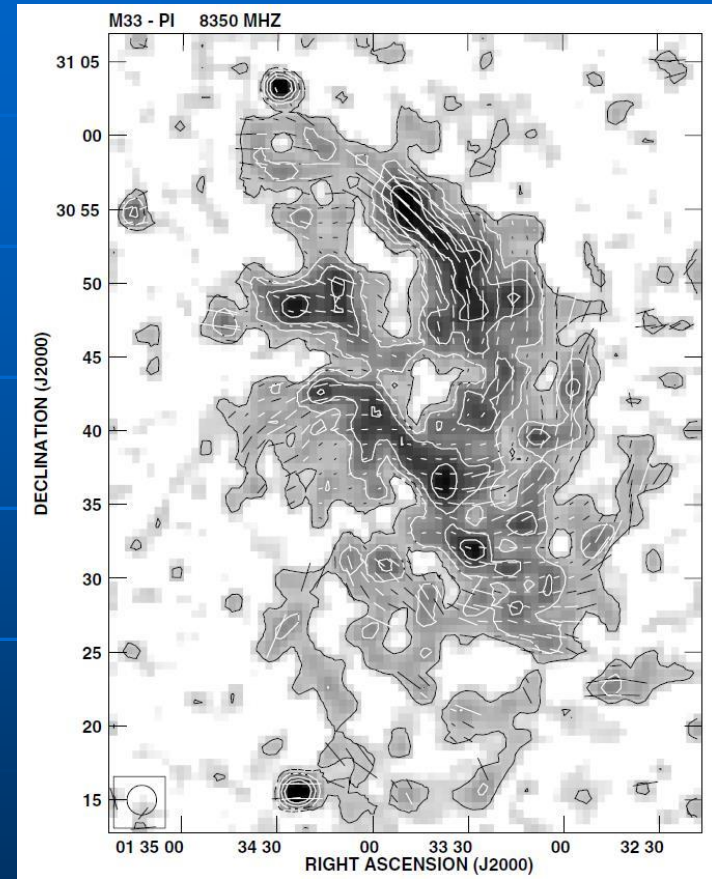


Q & U:
 $\sigma \approx 0.07 \text{ mJy} / 3' \text{ beam}$

Effelsberg 8.35 GHz (3.6 cm), BW = 1100 MHz



Total intensity:
 $\sigma \approx 0.5 \text{ mJy} / 2' \text{ beam}$



Q & U:
 $\sigma \approx 0.04 \text{ mJy} / 2' \text{ beam}$

Effelsberg is the world's most sensitive radio telescope to map diffuse polarized emission

- $\sigma \approx 0.07$ mJy/ 3' beam (Q&U at 4.85 GHz, 29h on-source to map a field of ≈ 1 deg²)
 - This corresponds to a surface brightness of diffuse emission of 0.5 μ Jy/ 15" beam
 - JVLA C-band D-array (3 GHz bandwidth): more than 100h on-source needed, FoV only ≈ 0.02 deg²
 - JVLA cannot see large-scale emission
- Effelsberg is more sensitive to diffuse emission than the JVLA

Present-day limits to measure magnetic field strengths with the Effelsberg telescope

Lowest detectable total intensity: $S_{\min} \approx 1.5 \text{ mJy}/3' \text{ beam}$
(confusion limited):

Equipartition strength of the total field, 1000 pc pathlength,
spectral index -0.8: $\approx 6 \mu\text{G}$

100 pc pathlength: $\approx 11 \mu\text{G}$

Lowest detectable polarized intensity: $S_{\min} \approx 0.2 \text{ mJy}/3' \text{ beam}$
(NOT confusion limited):

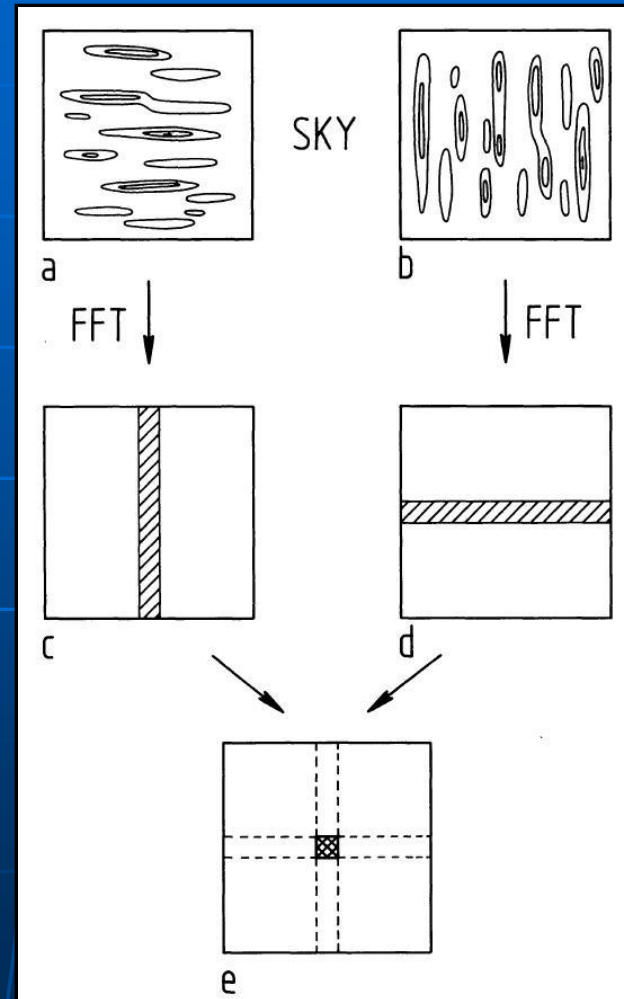
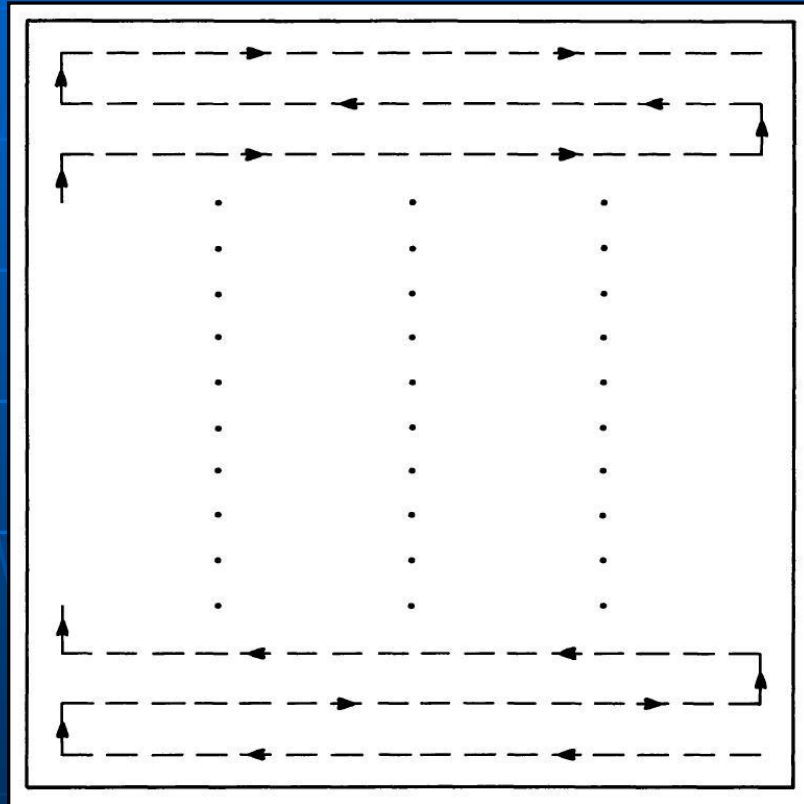
Equipartition strength of the ordered field, 1000 pc pathlength,
spectral index -0.8: $\approx 2.5 \mu\text{G}$

100 pc pathlength: $\approx 4.5 \mu\text{G}$

Reasons for the success of the Effelsberg systems

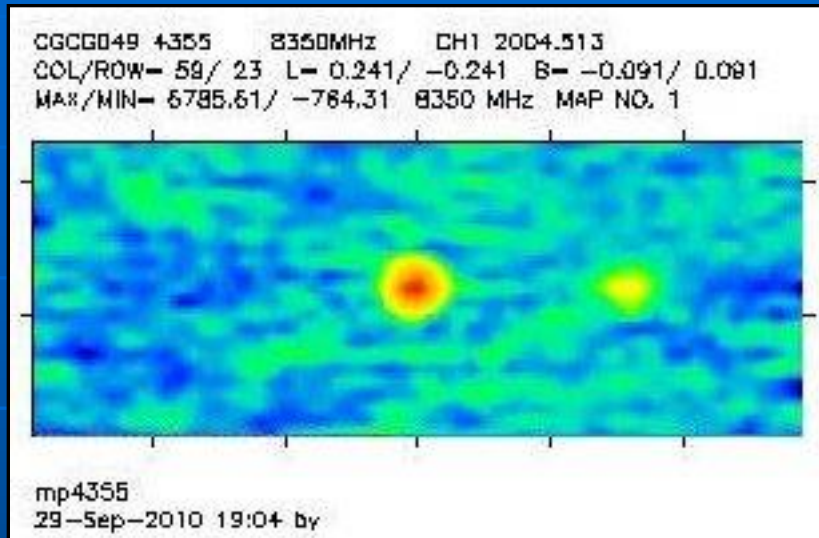
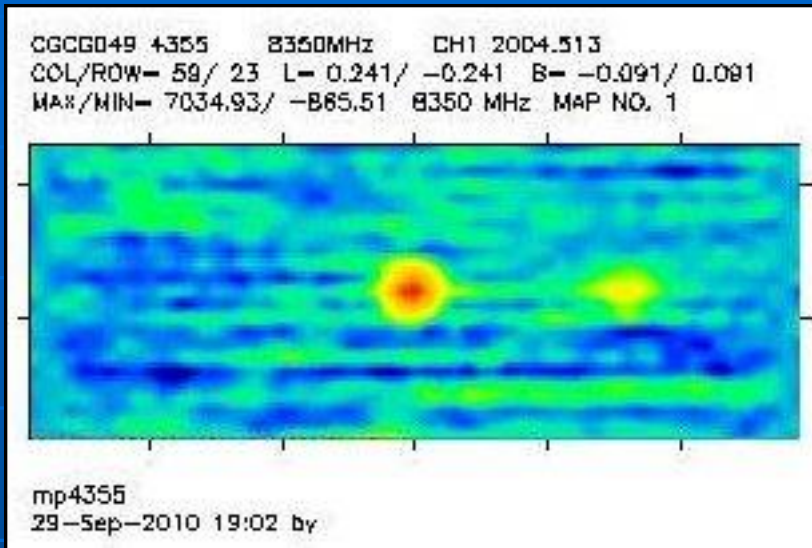
- Sensitivity (low system temperatures)
- Low degree of instrumental polarization
- Absolute gain stability by fast switching of the cal signal
- Relative gain stability between R and L
- Phase stability between Q and U
- Removal of quadratic terms of the correlator by using a 180° phase shift
- Dedicated software to reduce scanning effects

Suppressing scanning noise: Scanning in different directions

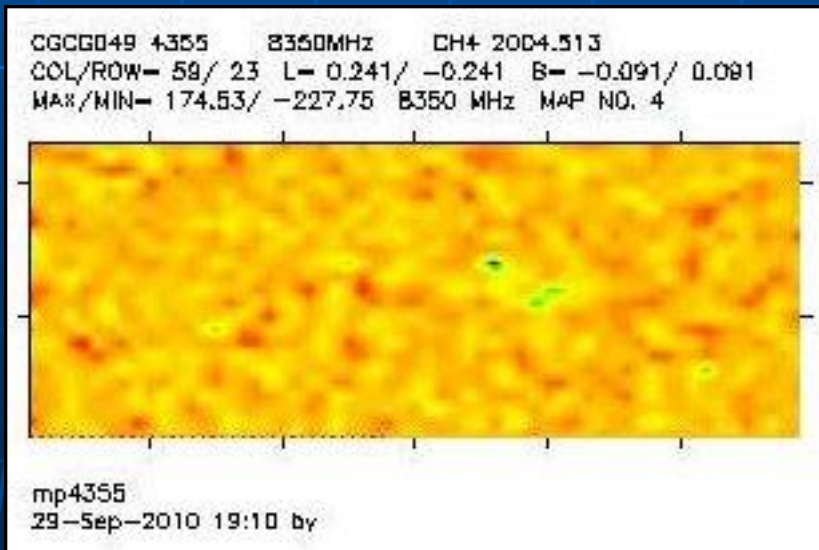
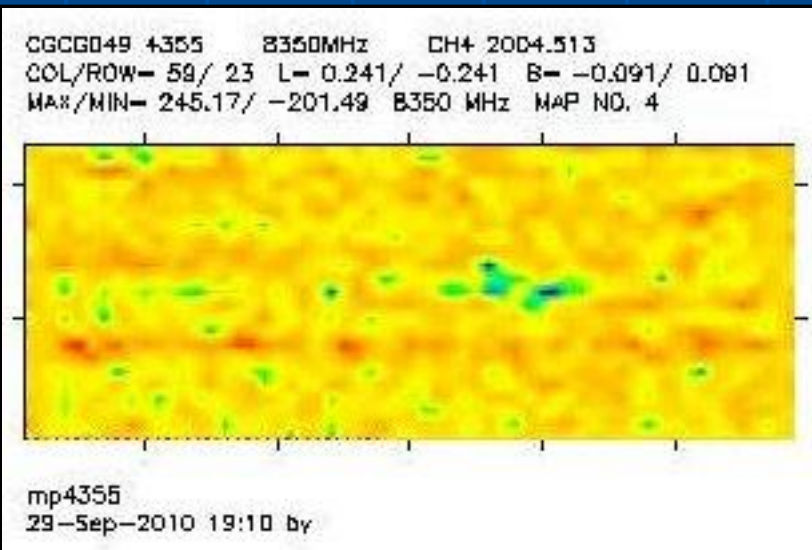


Suppressing scanning noise (8.35 GHz)

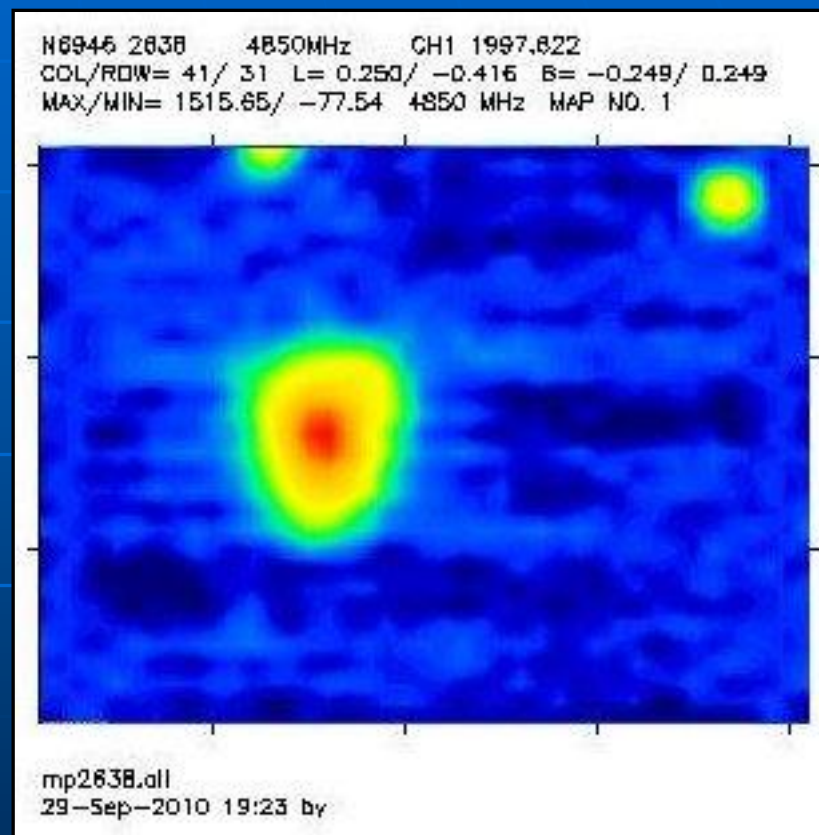
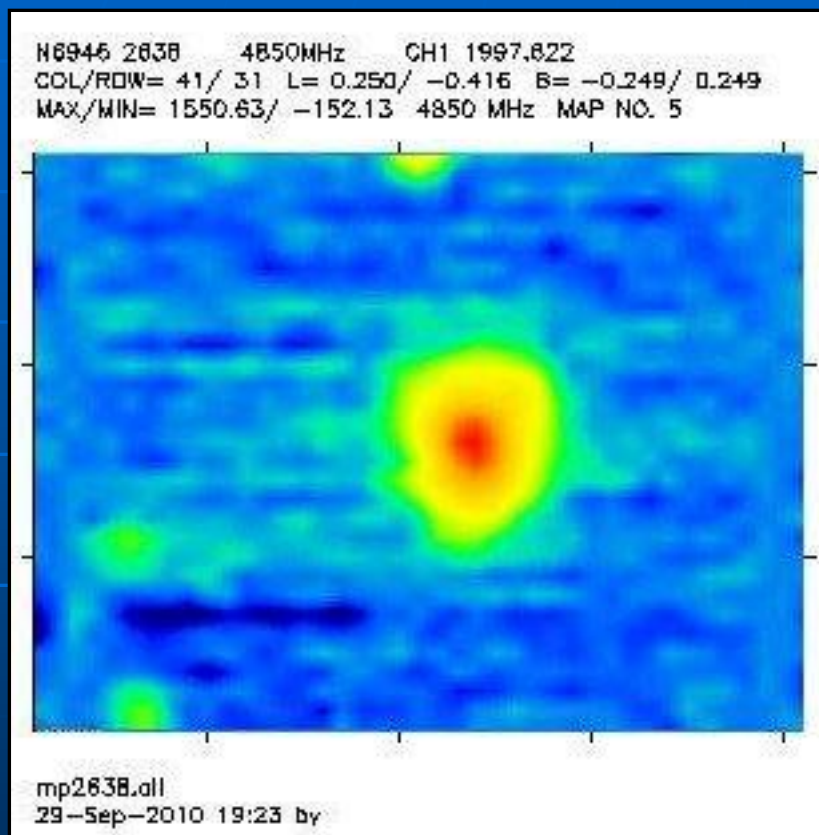
I



Q



Suppressing scanning noise: Dual-horn observations (4.8 GHz)



Continuum & polarization mapping: the future

- New wide-band receivers allow even higher sensitivity
- Transformation from linear to circular polarization is crucial, but hardware hybrids cannot be used in the new wide-band systems
→ **Digital conversion** (see e.g. Das et al. 2011)
- **Switched calibration signal** is crucial for stability (but blank / sync information needed)
- **Digital correlation** of many frequency channels for two signal inputs (R and L)
- **Multi-channel polarimetry** allows to apply *RM* Synthesis and to compute *Faraday spectra*)
- **Improved software** for scanning removal, cleaning, etc: *NOD3*